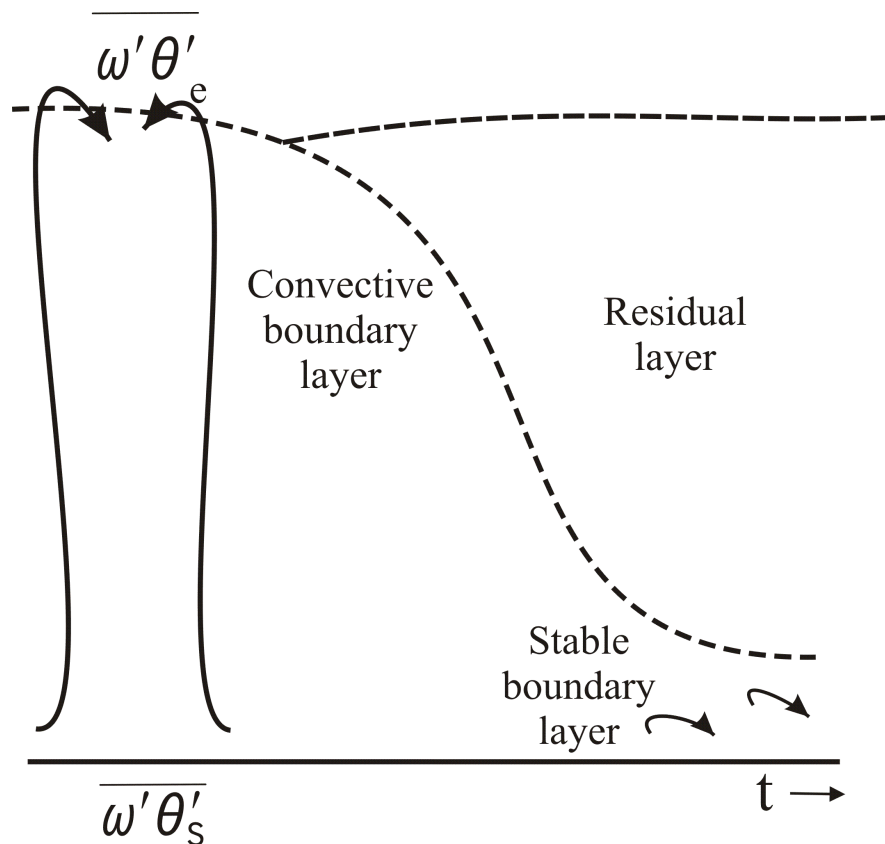


# Vertical and temporal evolution of turbulence spectra in the late afternoon transition

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Fabienne Lohou, Marie Lothon, David Pino, Jordi Vilà  
and BLLASTers

Improve our understanding of the vertical structure of the BL during the afternoon transition



Is there a decoupling between surface and higher levels during LAT?

- **Spectral analysis** at different heights (aircraft, surface measurements, LES)

- Evolution of the characteristic length scales of  $w$

- Evolution of the shape of the turbulence spectra

Until now :

## **Shape of the spectra**

-Many models for convective conditions

:

                  Kaimal et al. (1976) (validated  
in surface: Kansas, Minnesota  
experiments), von Kàrmàn (1948)

- Analytical spectra models not always  
adapted within the entire CBL (Lothon  
et al. (2009))

## Kristensen et al. (1989) general kinematic spectral model

**Hypothesis** : Anisotropic horizontally homogeneous vertical velocity field

$$\frac{S(k)}{\sigma_w^2} = \frac{l_w}{2\pi} \frac{1 + \frac{8}{3} \left(\frac{l_w k}{a(\mu)}\right)^{2\mu}}{\left(1 + \left(\frac{l_w k}{a(\mu)}\right)^{2\mu}\right)^{\frac{5}{6\mu} + 1}}$$

$$a(\mu) = \pi \frac{\mu \Gamma\left(\frac{5}{6\mu}\right)}{\Gamma\left(\frac{1}{2\mu}\right) \Gamma\left(\frac{1}{3\mu}\right)}$$

$$\lambda_w = \left\{ \frac{5}{3} \sqrt{\mu^2 + \frac{6}{5}\mu + 1} - \left(\frac{5}{3}\mu + 1\right) \right\}^{1/(2\mu)} \frac{2\pi}{a(\mu)} l_w$$

**Two characteristics** :

$l_w$  (integral scale)

$\mu$  (shape)

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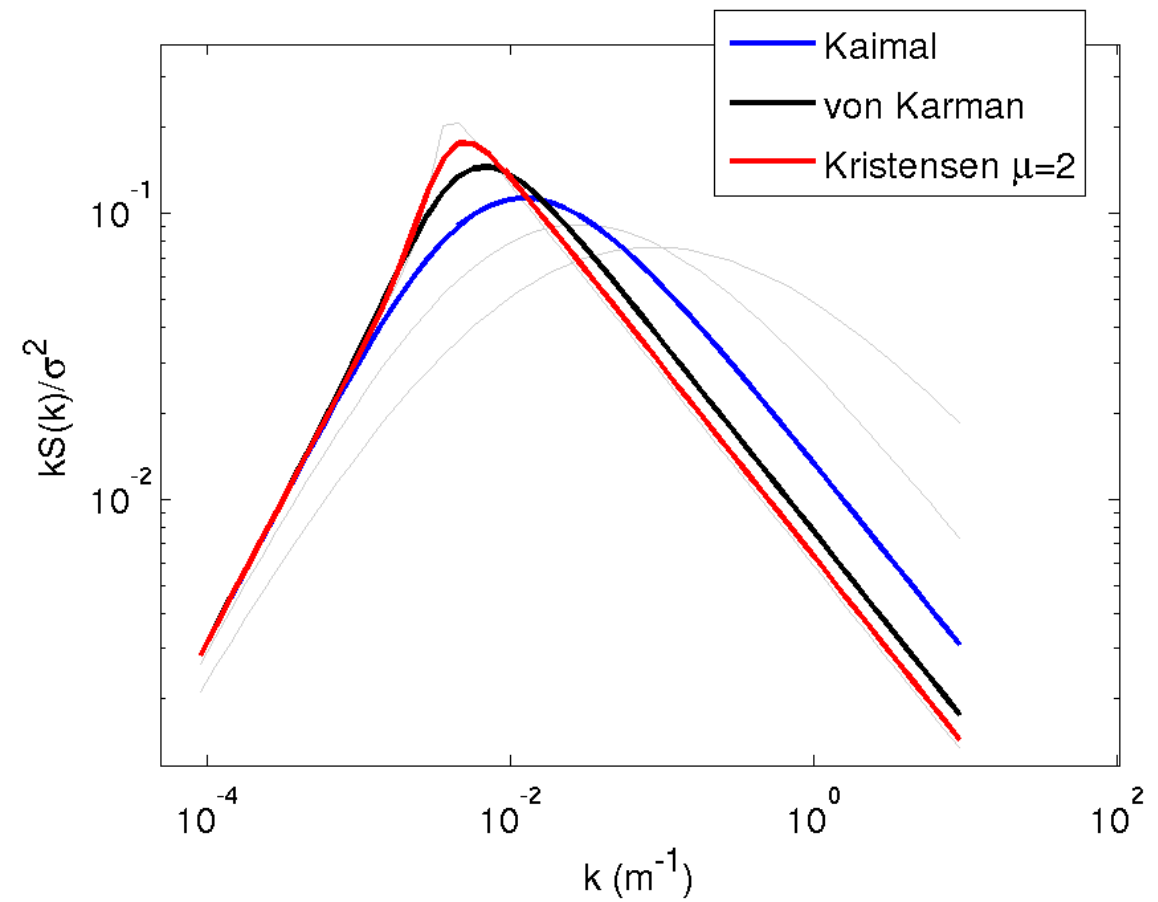
    → Kristensen et al. analytical model (shape parameter)

- Few studies on its temporal evolution

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## Lengthscales

-Many ways to define those scales

→ **wavelength of the energy**

**spectrum peak** (energy production) :

*Nieuwstadt and Brost (1986)*

*Grant (1997)*

→ **integral scale** (energy-  
containing eddies): *Sorbjan (1996)*

→ **weighted integral of the**

**spectrum** : *Pino et al. (2006)*

→ etc...

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## During LAT:

*constant*

*decrease*

*increase*

*increase, except for  $w$  (constant)*



## What is our contribution?

Until now :

### Lengthscales

- Many ways to define those scales
  - **wavelength of the energy spectrum peak** (energy production) :  
*Nieuwstadt and Brost (1986)*  
*Grant (1997)*
  - **integral scale** (energy-containing eddies): *Sorbjan (1996)*
  - **weighted integral of the spectrum** : *Pino et al. (2006)*
  - etc...
- Lack of agreement in the evolution of those scales during LAT

- **Spectral analysis during the LAT (BLLAST + LES) using Kristensen et al. analytical spectra model**
  - 1/ Evolution of the characteristic length scales of  $w$  (2 methods) during the LAT
  - 2/ Evolution of the shape of the turbulence spectra



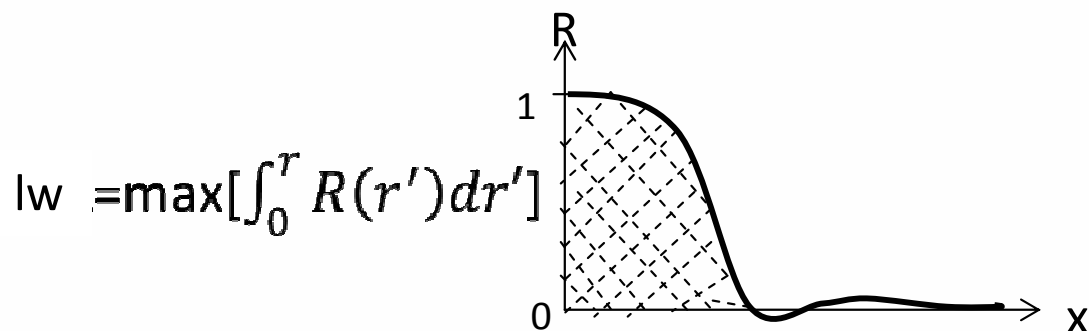
1) Case study : 20 June 2011

2) Use of an **analytical spectra model** for convective BL as a reference (Kristensen et al.,1989)

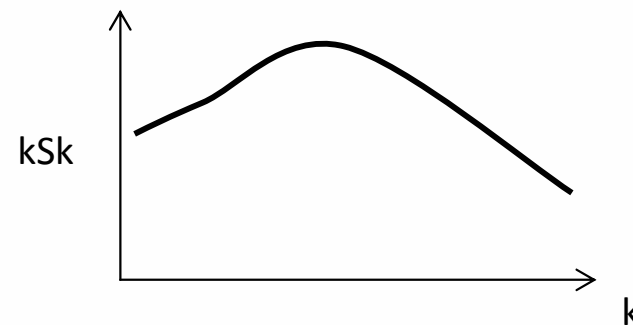
- Calculation of the turbulence spectra (with LES and BLLAST data) every 30min :  
compromise between sufficient number of eddies and stationary conditions
- Search of  $(\mu, l_w)$  which gives the best fit

3) **Calculation of the integral length scales** from 2 methods :

(1) with the autocorrelation function



(2) from the analytical spectra model ( $l_w$ )



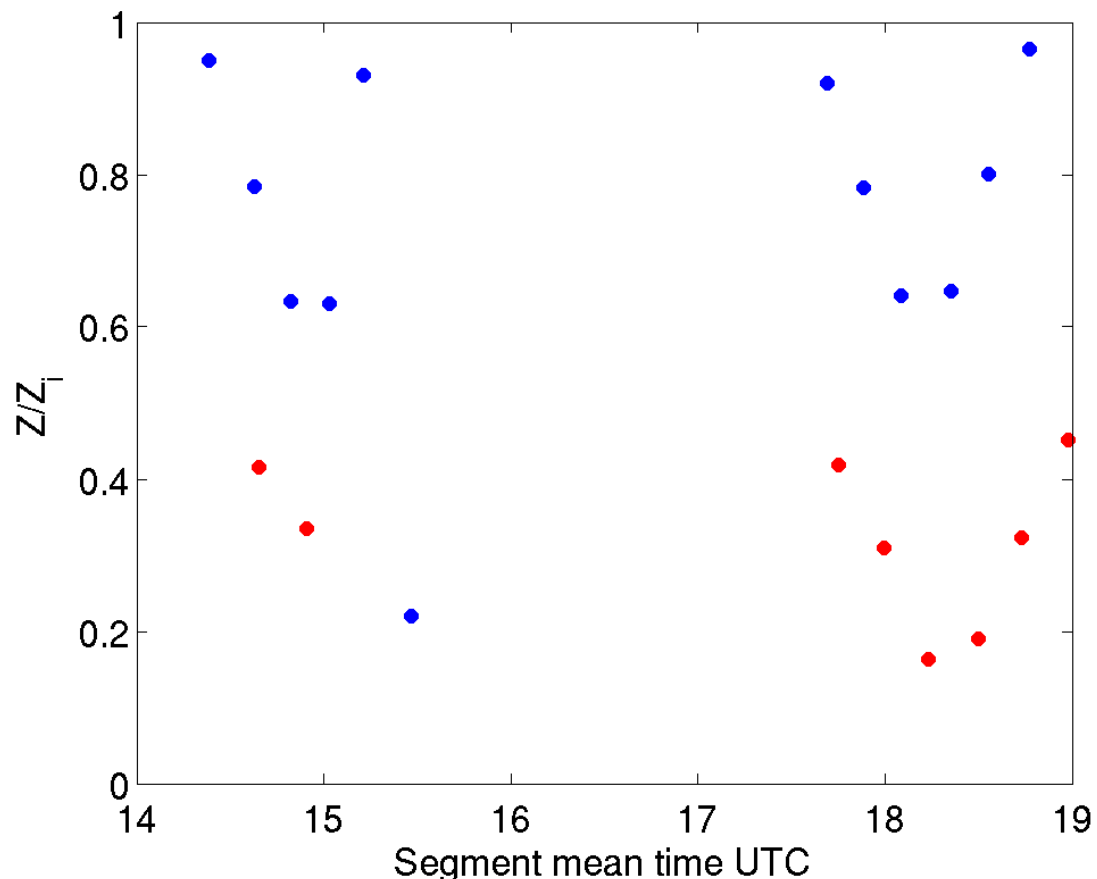
# Data used to calculate the spectra

- BLLAST data : Various surfaces, 60m mast, aircrafts
- LES

## Piper Aztec and Sky Arrow on 20 June

2 flights for each in the afternoon

*Three // legs, 6 heights, 3 latitudes*

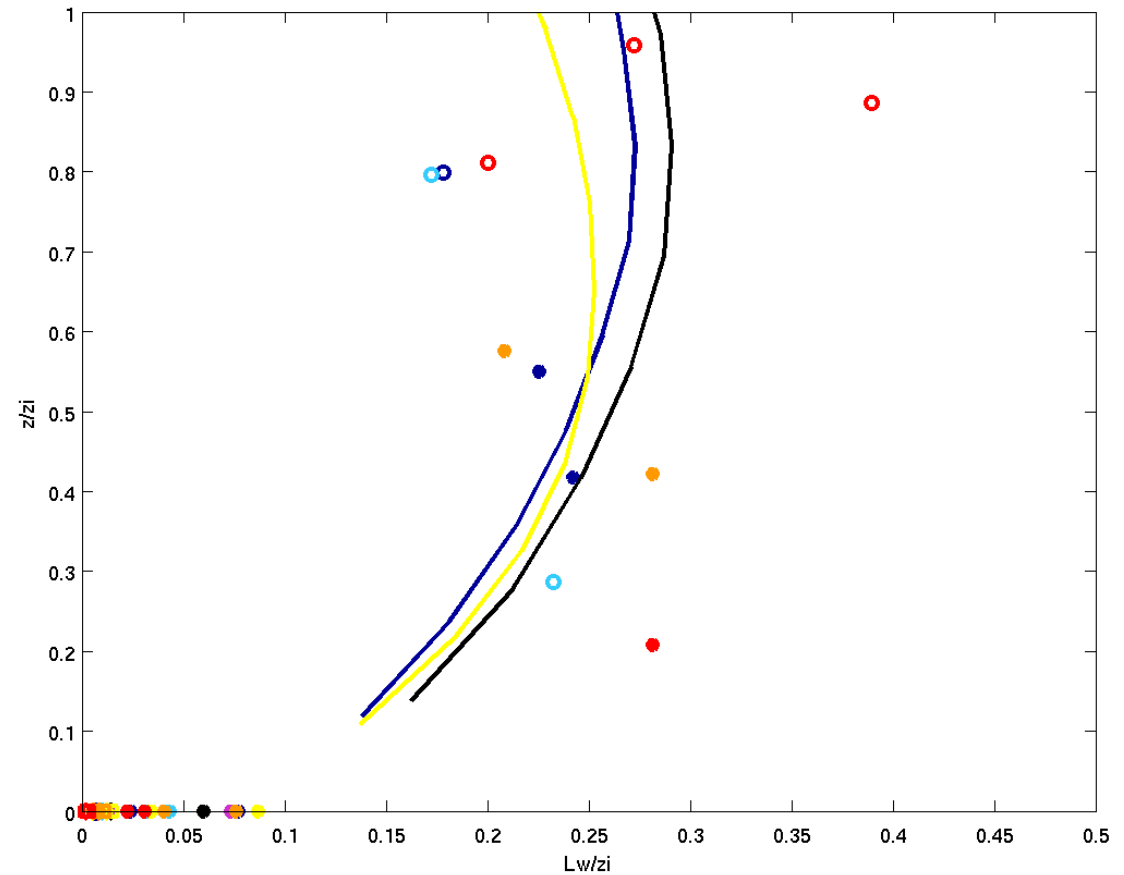
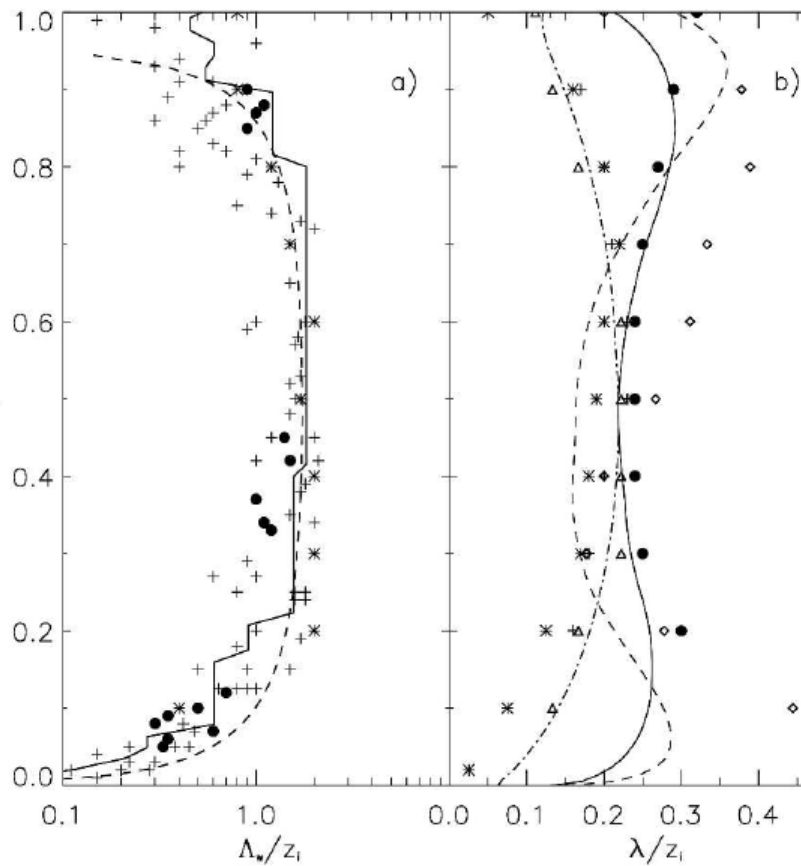


- Aircrafts: dense observations of the turbulence decay within the ML
- Measurements over various surfaces

# Validation of results found in literature for vertical profiles of lengthscales at 12h

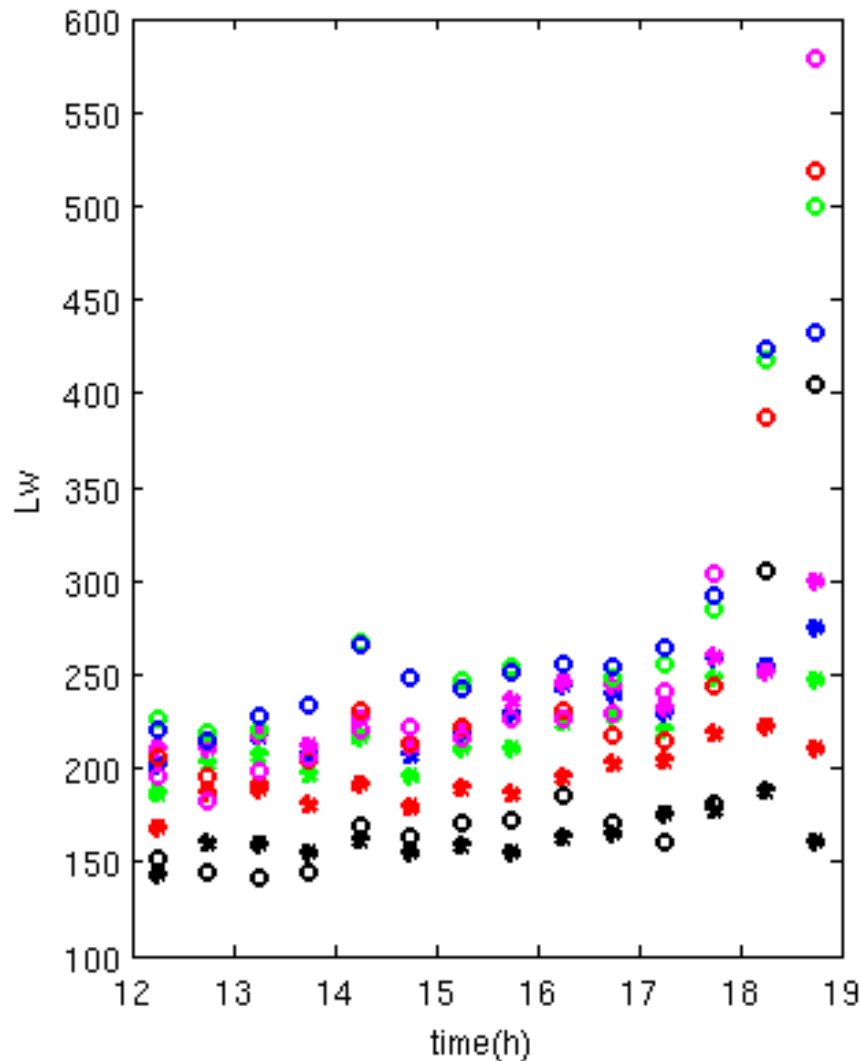
Dosio et al. (2005)

Vertical profiles of  $w$  lengthscales from BLLAST data (points), and LES (continuous lines) evolving with time (color changing from black to red)



➔ At 12h, same lengthscales than Dosio et al :  $lw \sim 0.25 z_i$

# Temporal evolution of $w$ integral length scales from the 2 methods at different heights. **With LES**

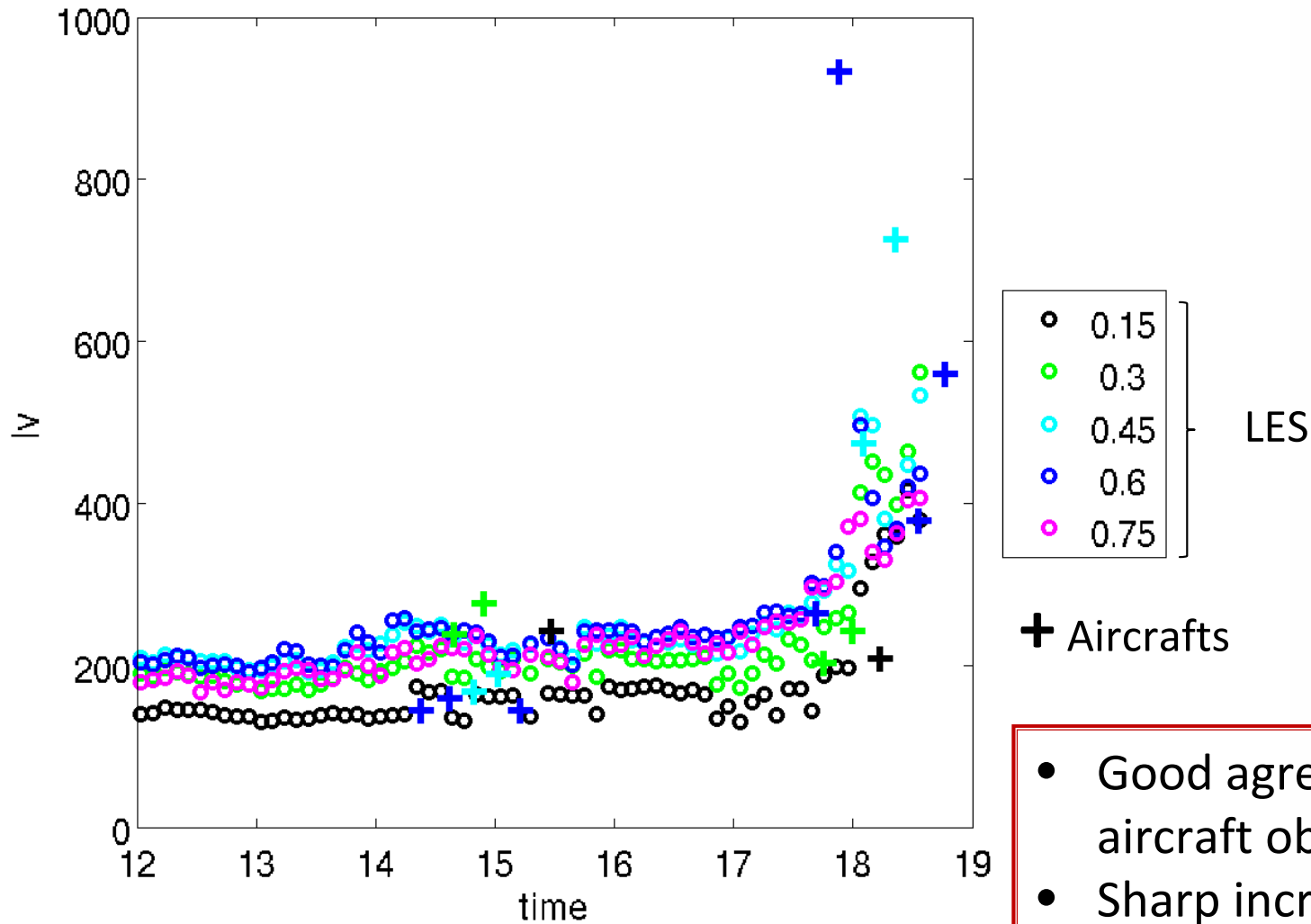


*From autocorrelation function*

*From analytical spectra model*

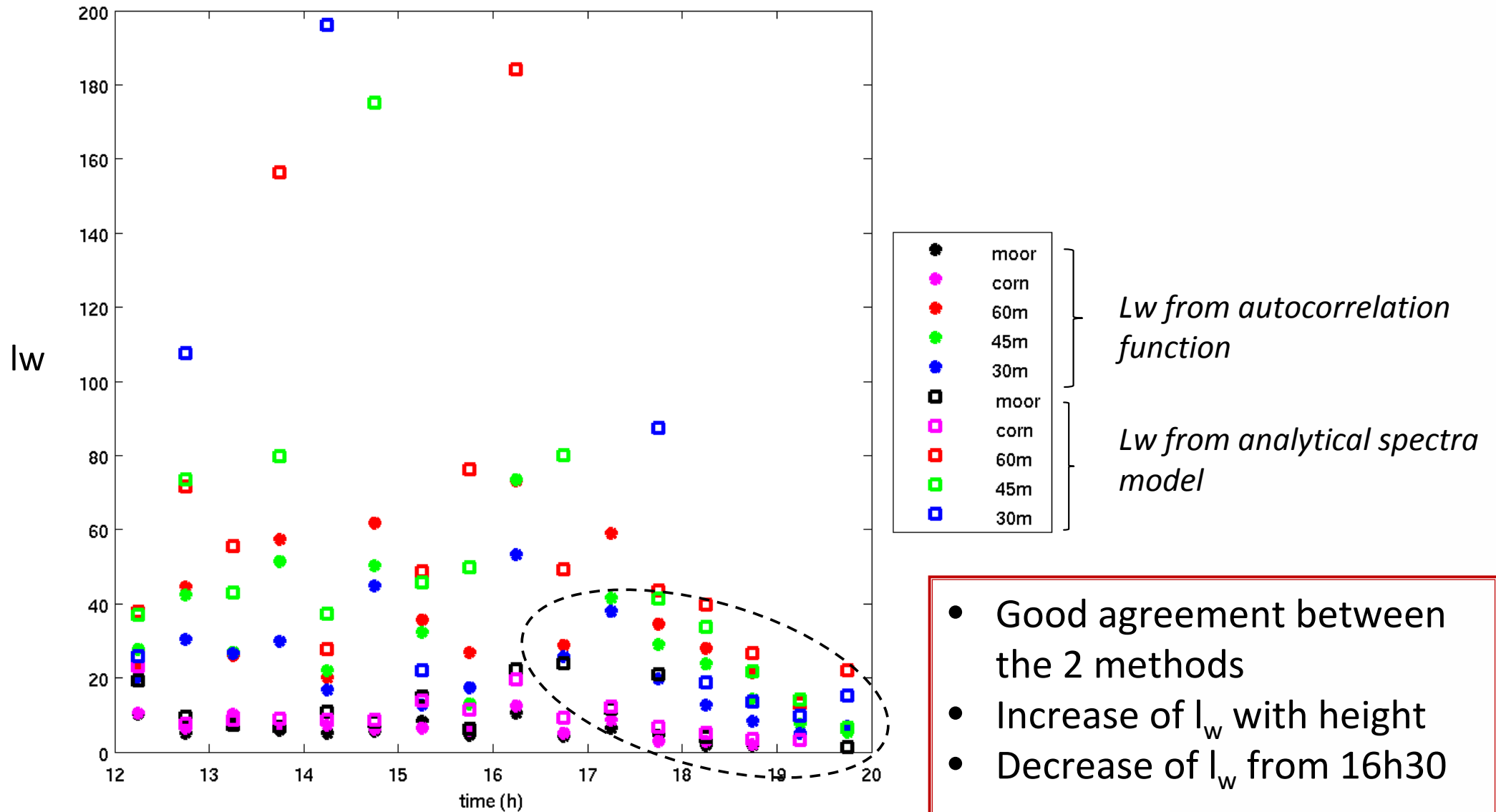
- 2 consistent methods
- Increase with altitude
- Both increase in the afternoon, and after 17pm with a different amplitude

### Evolution of the $I_w(z/z_i, t)$



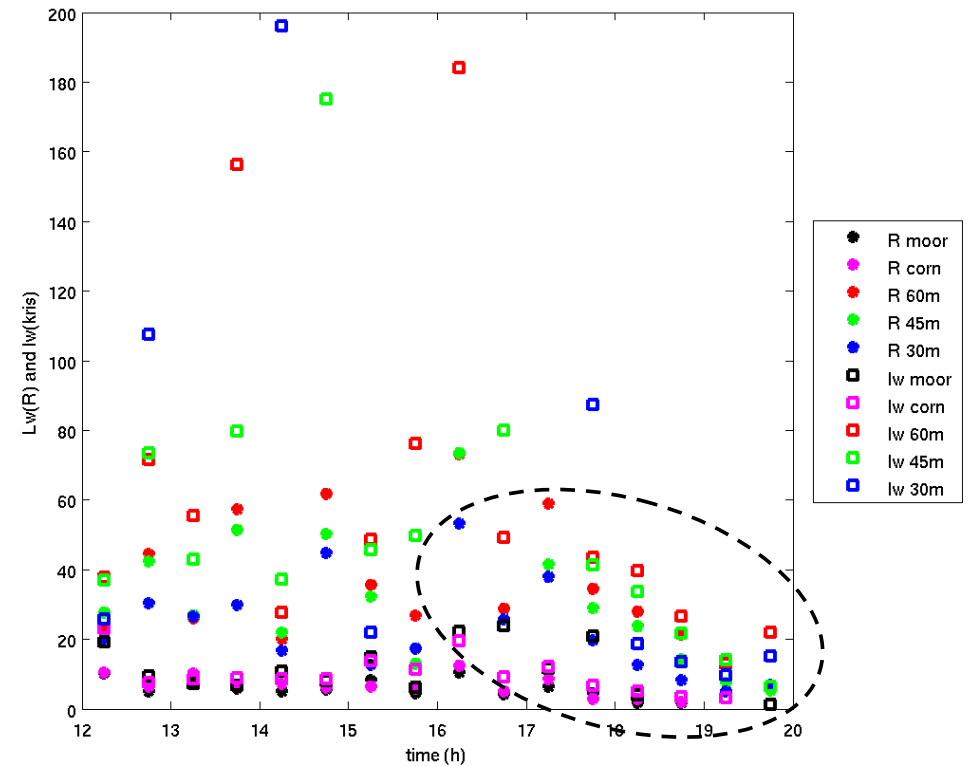
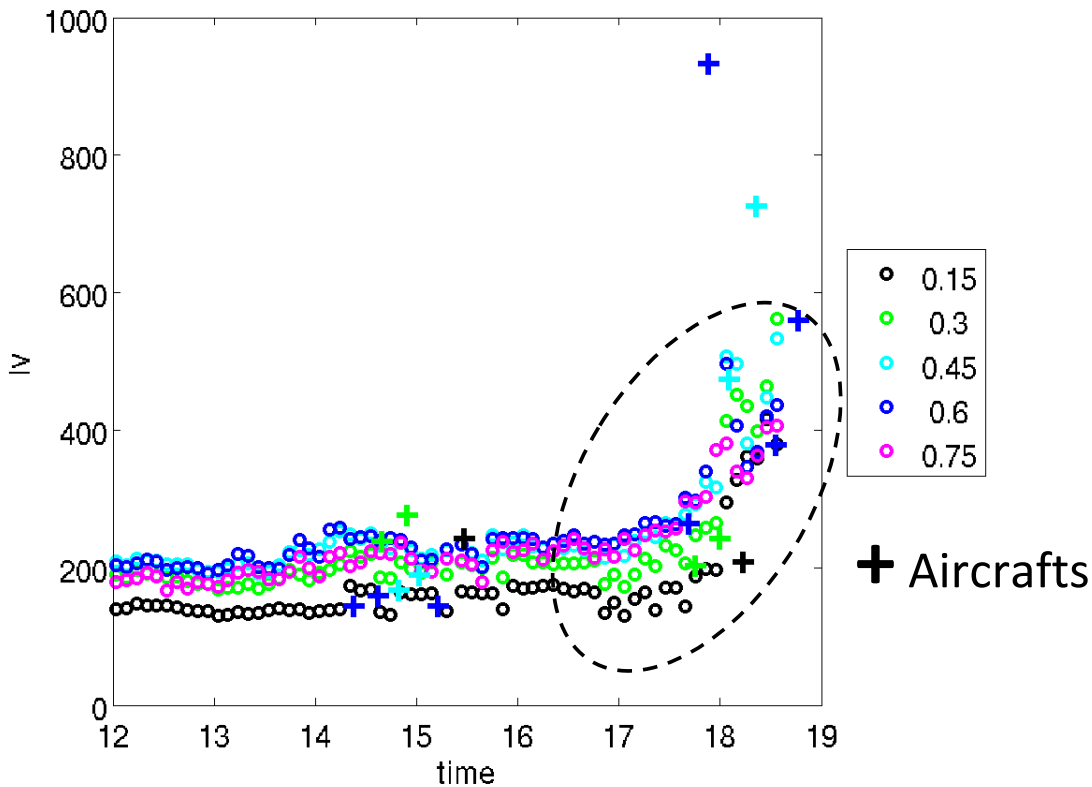
- Good agreement between aircraft observations and LES
- Sharp increase of  $I_w$  from 17h

# Temporal evolution of $w$ integral length scales from the 2 methods. Surface and 60m mast



- Good agreement between the 2 methods
- Increase of  $l_w$  with height
- Decrease of  $l_w$  from 16h30

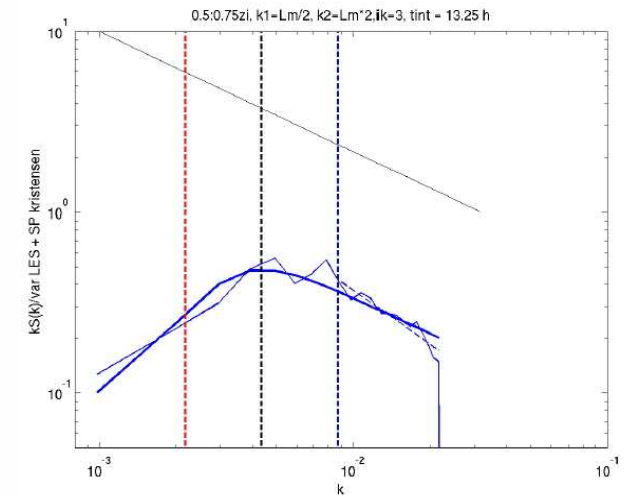
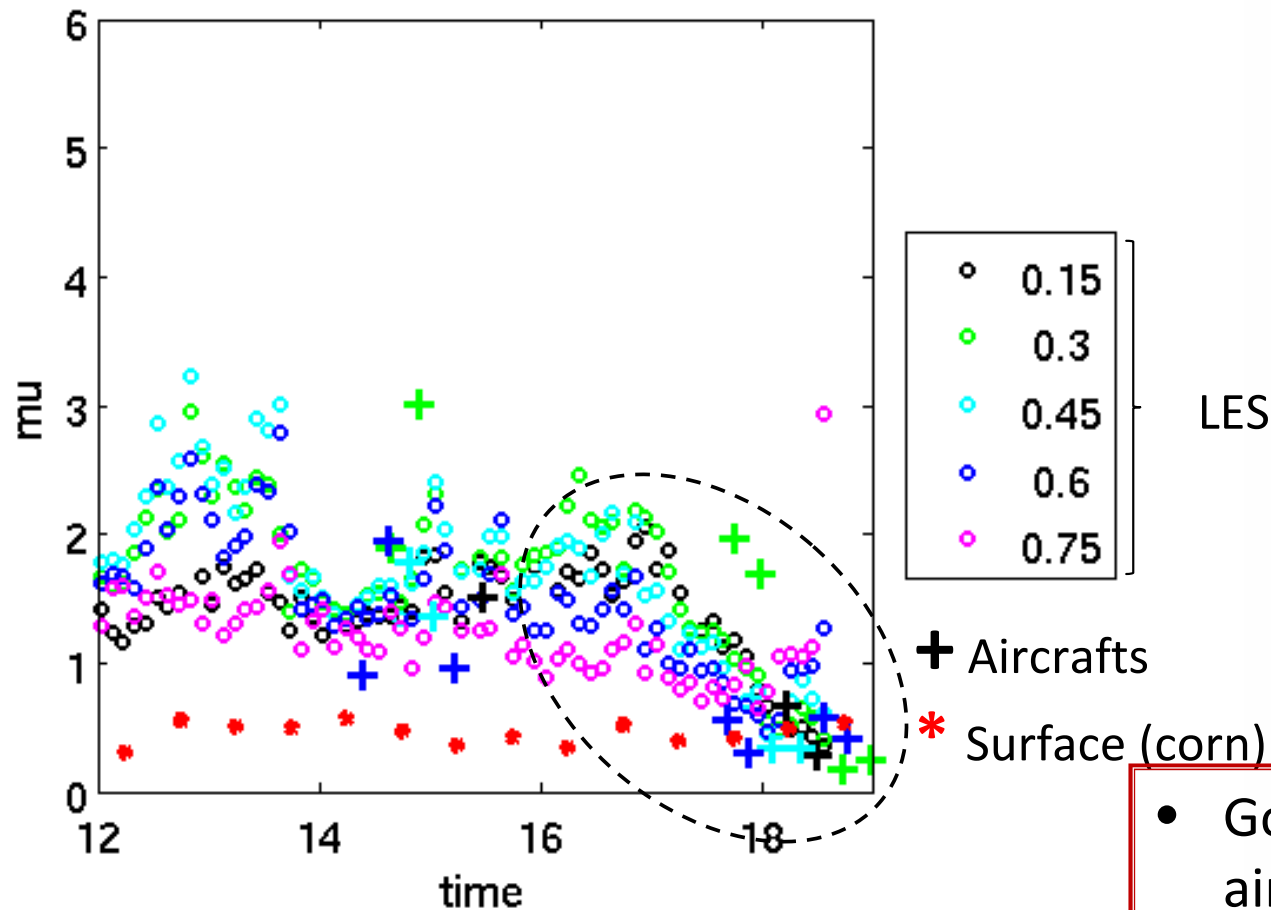
## Decoupling between surface layer and the layer above



- LES could fill the gap between surface and aircraft measurements
- How low can we estimate  $l_w$  with LES according the resolution ?



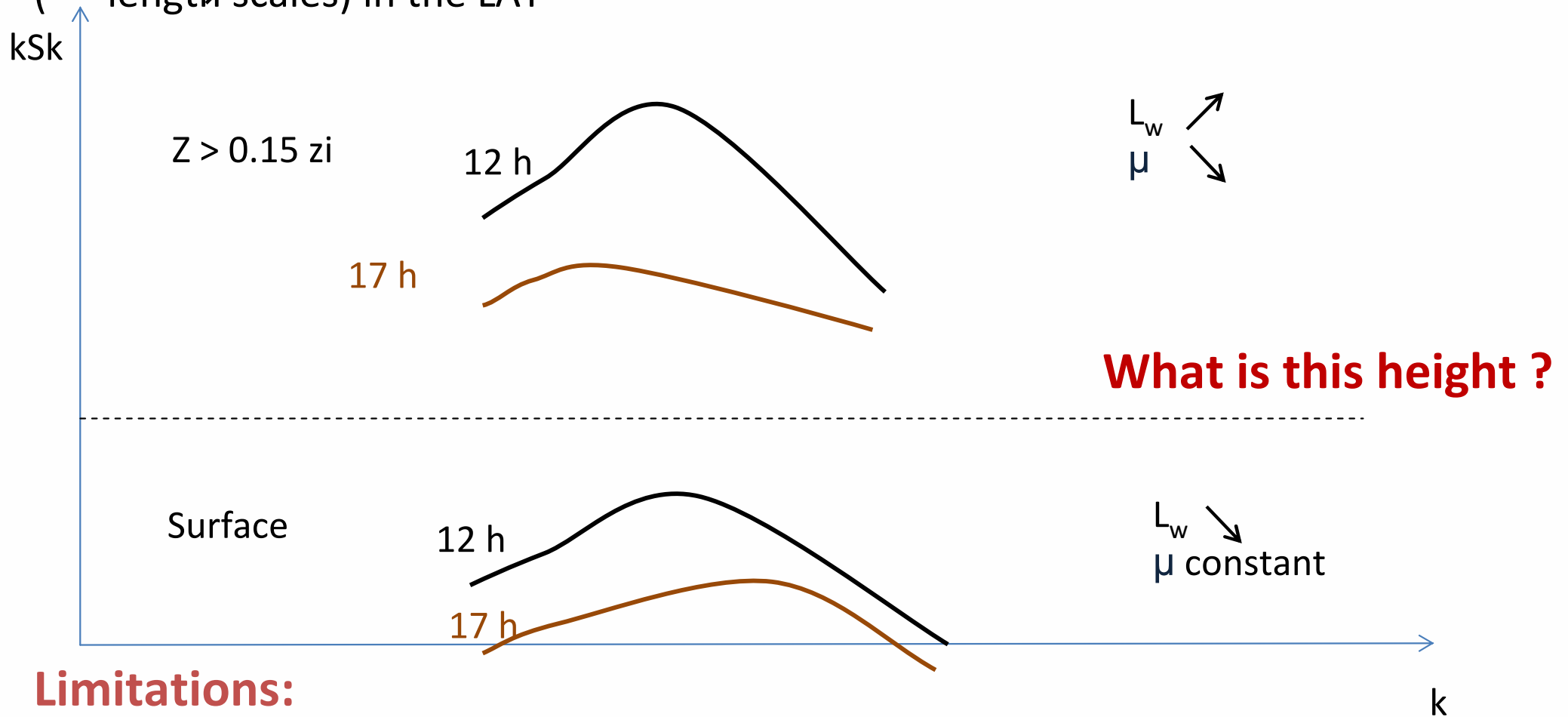
## Evolution of $\mu(z/z_i, t)$



- Good agreement between aircrafts and LES
- From 16h30, the spectra flatten above 0.15 $z_i$
- At surface,  $\mu$  remains constant (Kaimal !)

## Conclusions:

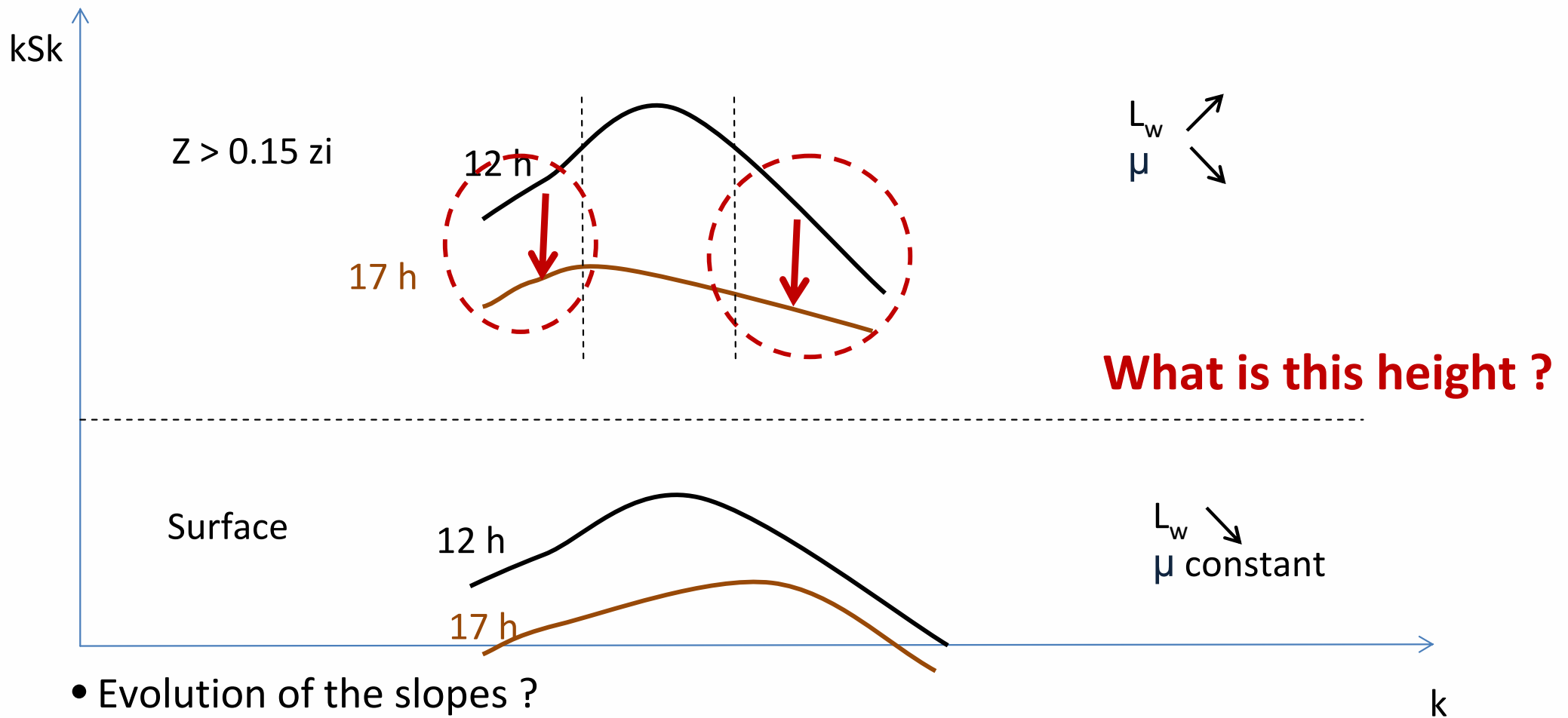
- Decoupling between the surface layer (  $\downarrow$  length scales) and the overlying layer (  $\uparrow$  length scales) in the LAT



## Limitations:

- Hypothesis : homogeneous isotropic turbulence
- Z<sub>i</sub> estimation

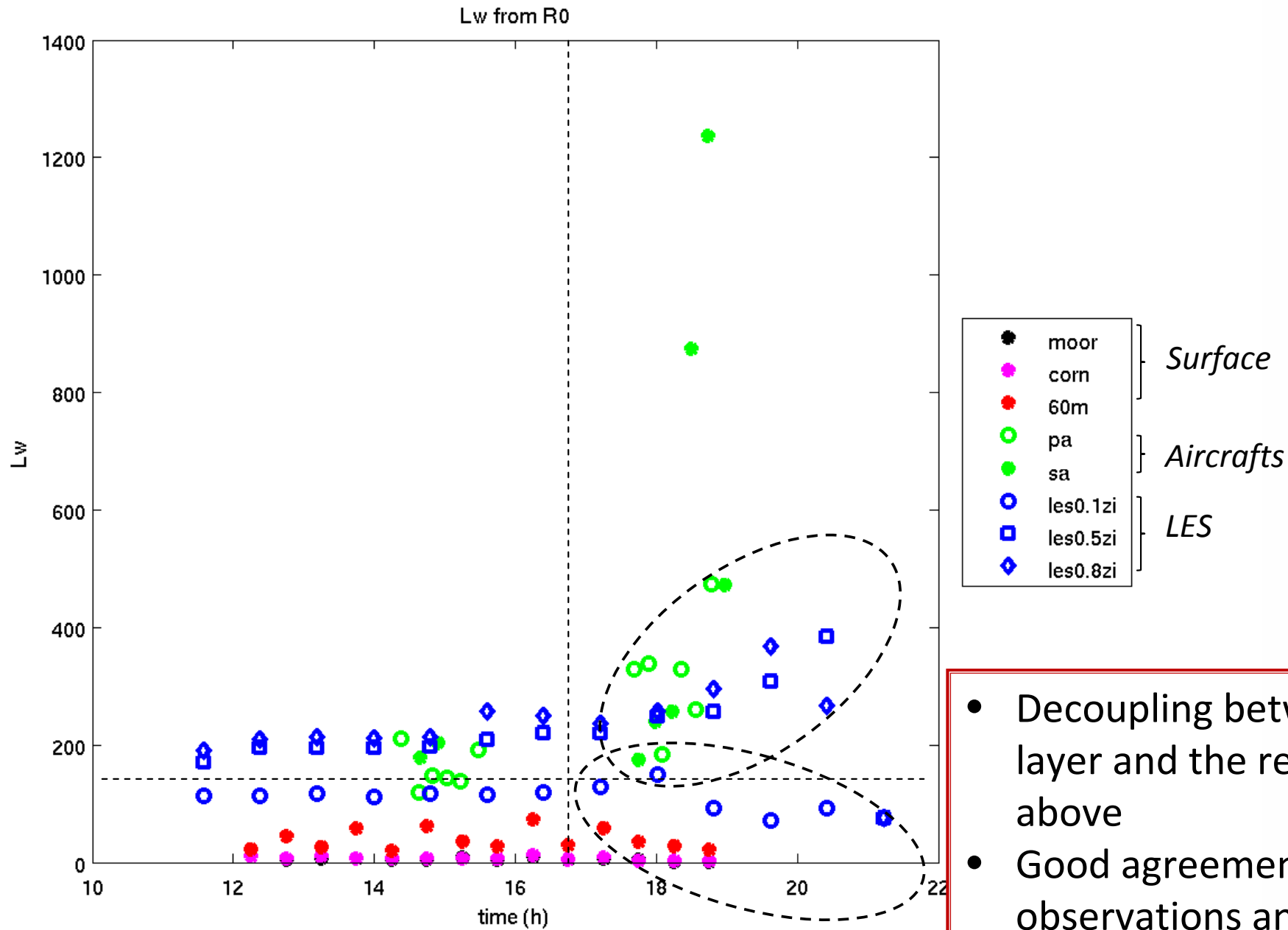
## Perspectives:



- Evolution of the slopes ?
- Evolution of the contribution of the different frequency domains ?
- Extend the study to other variables (T, u, v)

Thank you for your attention !

# Temporal evolution of $w$ integral length scales at different levels, from observations (surface, 60m, aircrafts) and LES



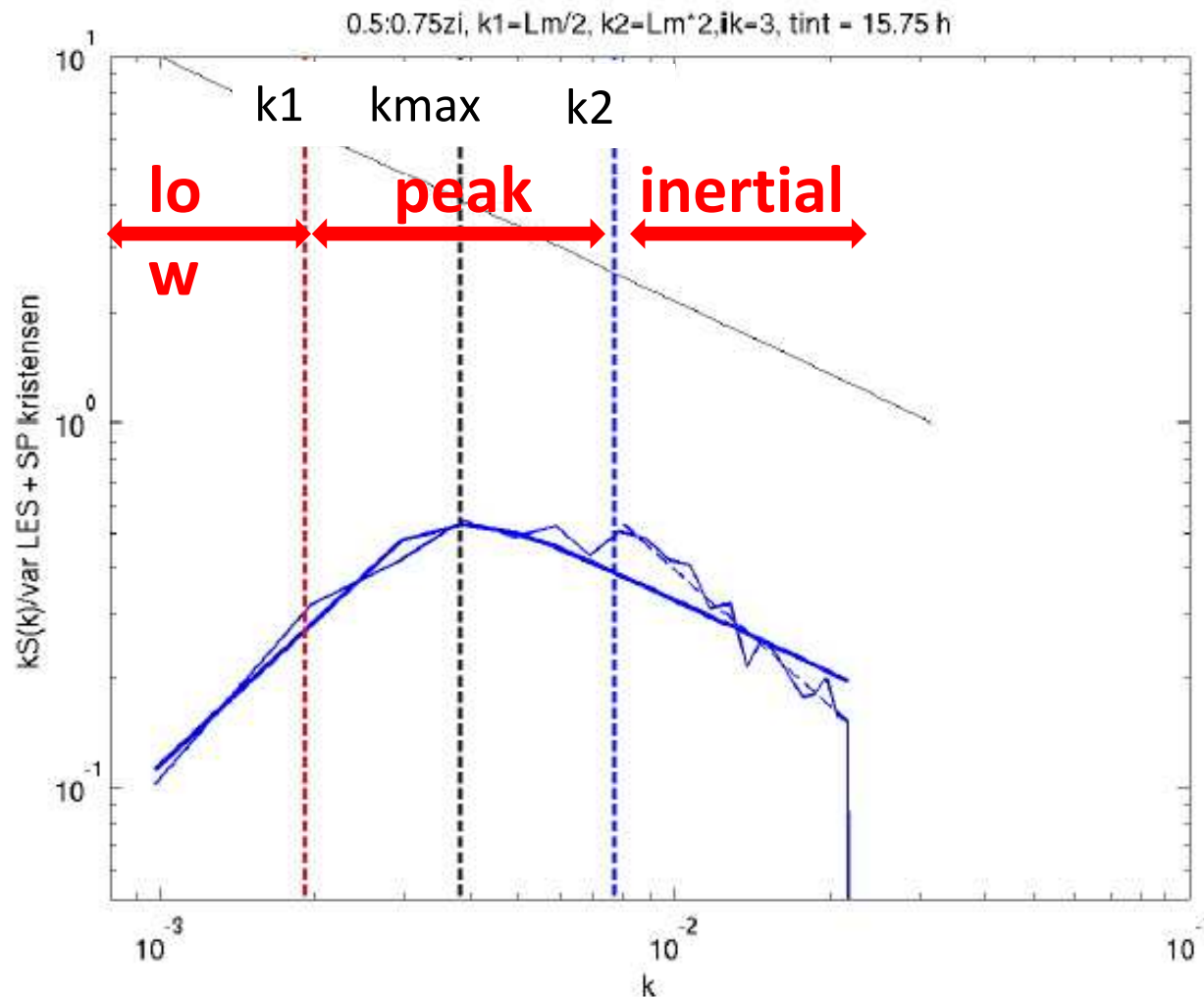
- Decoupling between surface layer and the residual layer above
- Good agreement between all observations and LES

# Spectral analysis

## Three domains defined

Limits are  $\mu$ -depending in case of observations, fixed when applied to shorter LES-range.

1. The low wavenumber range
2. Around the peak of the maximum of the spectra
3. The inertial subrange



OBS

$$k_1 = 5/\mu k_{max}$$

$$k_2 = \mu/5 k_{max}$$

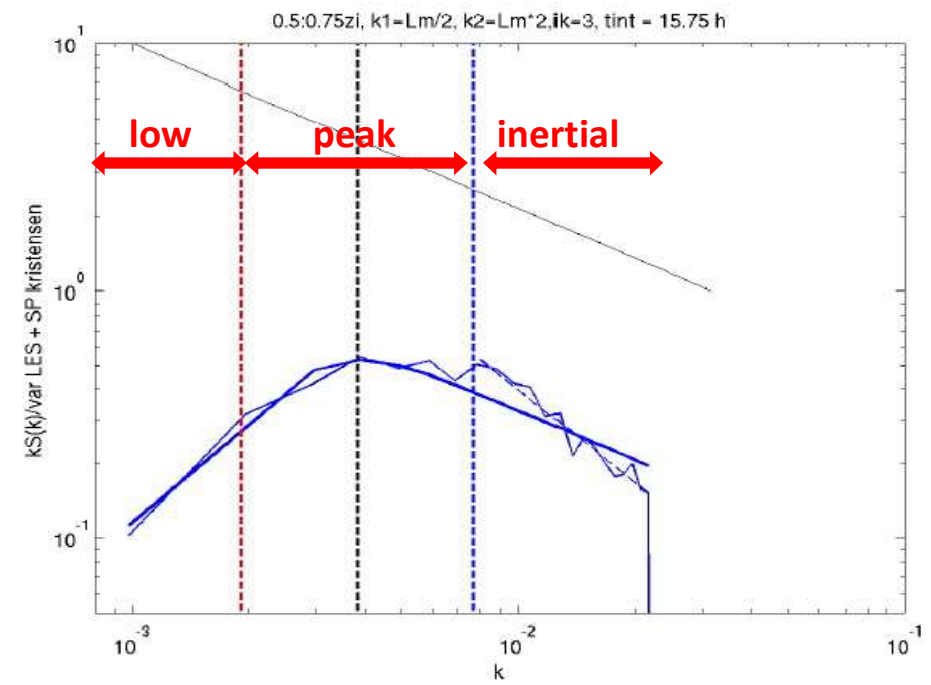
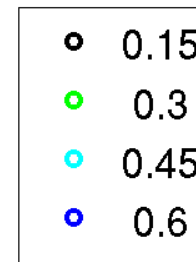
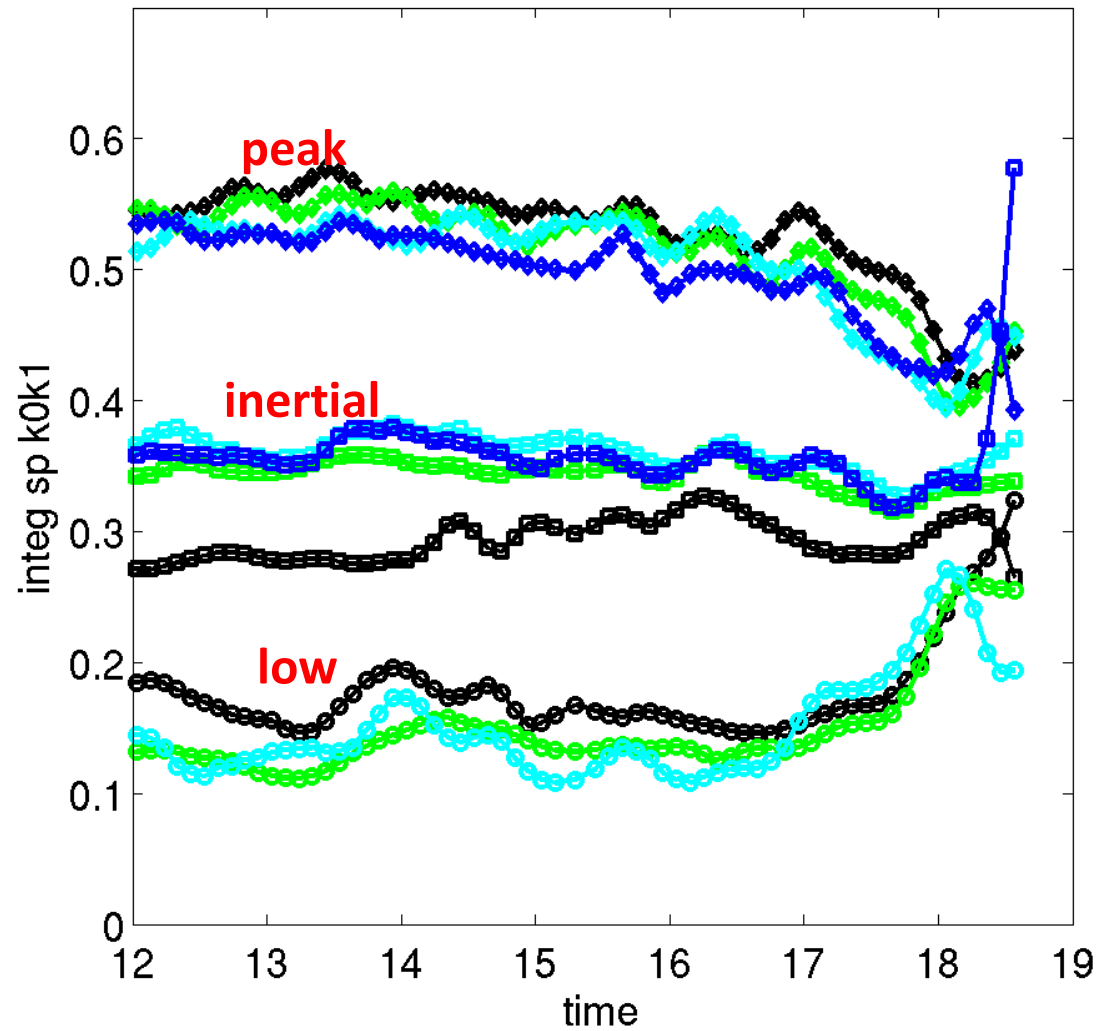
LES:

$$k_1 = \frac{1}{2} k_{max}$$

$$k_2 = 2 k_{max}$$

# Contribution (%) of the 3 frequency domains

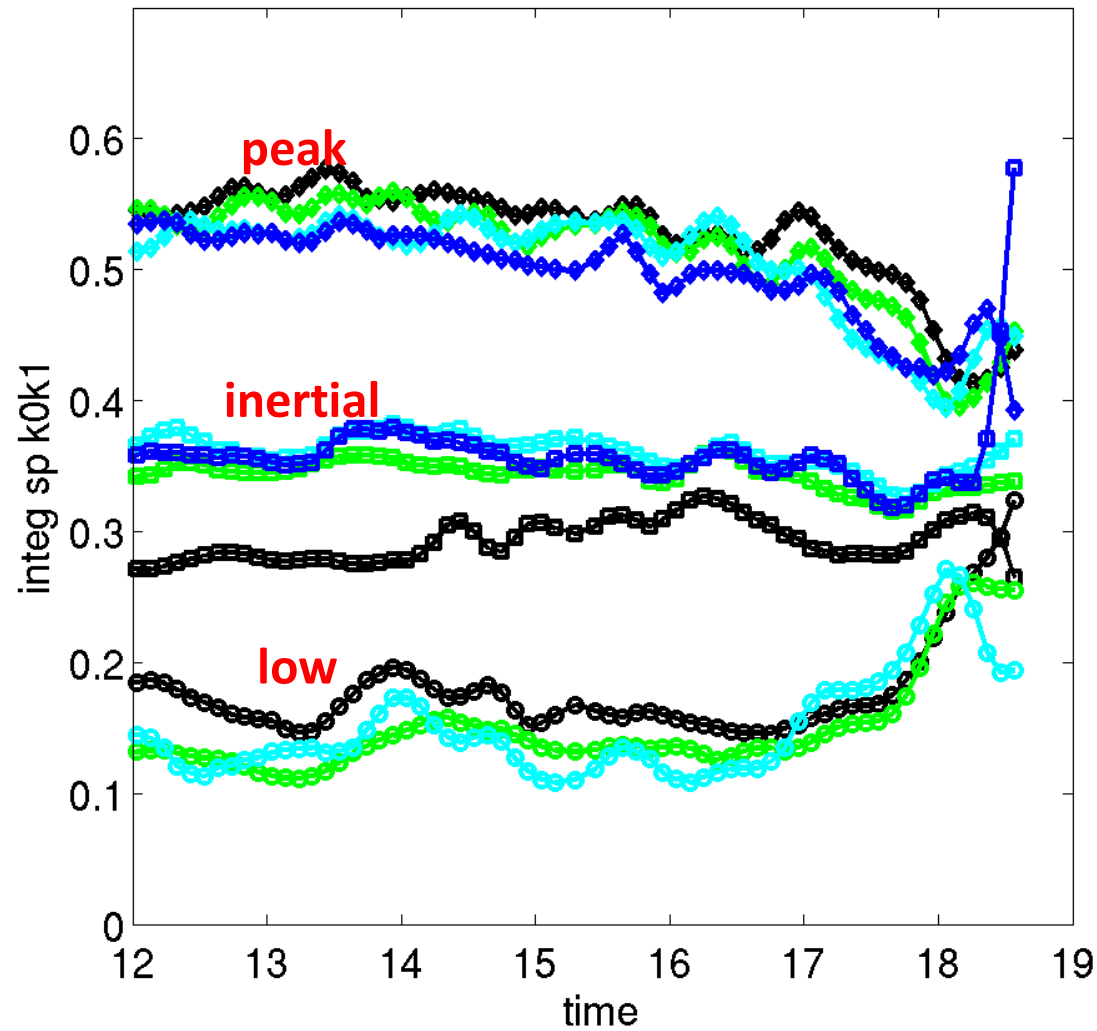
evolution of the integral of the spectra in the 3 domains (o d s)



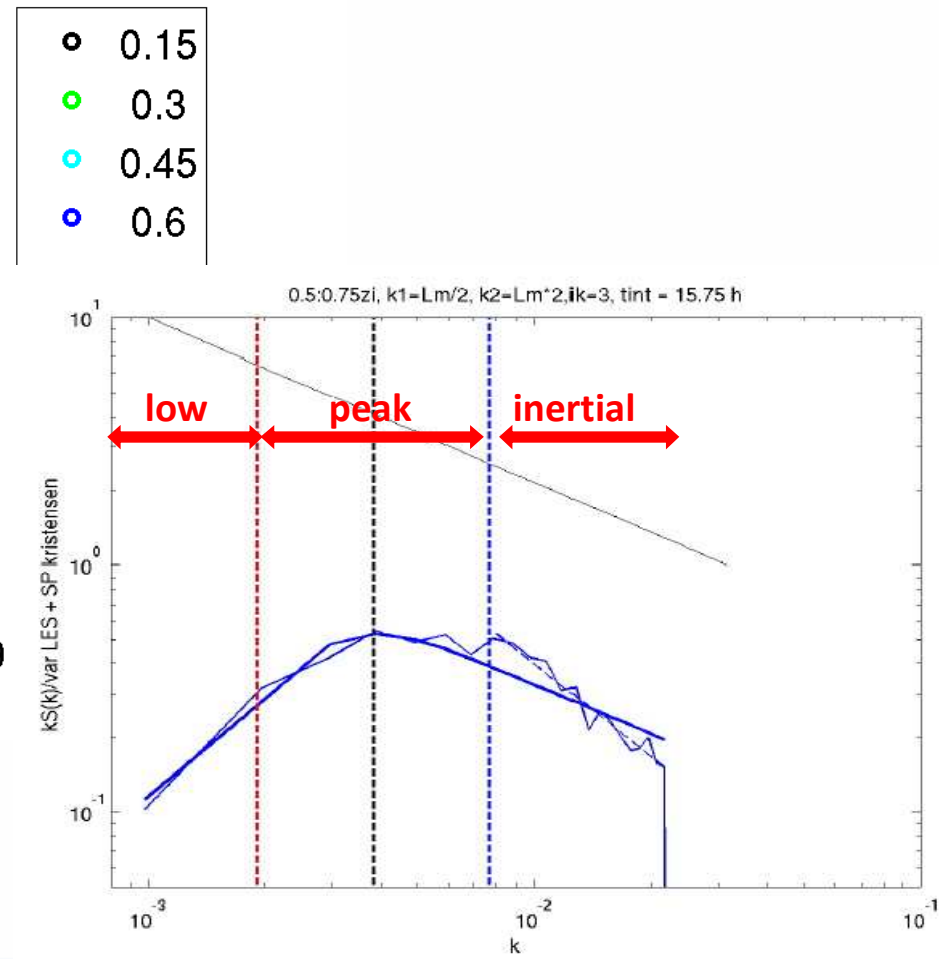


# Contribution (%) of the 3 frequency domains

evolution of the integral of the spectra in the 3 domains (o d s)



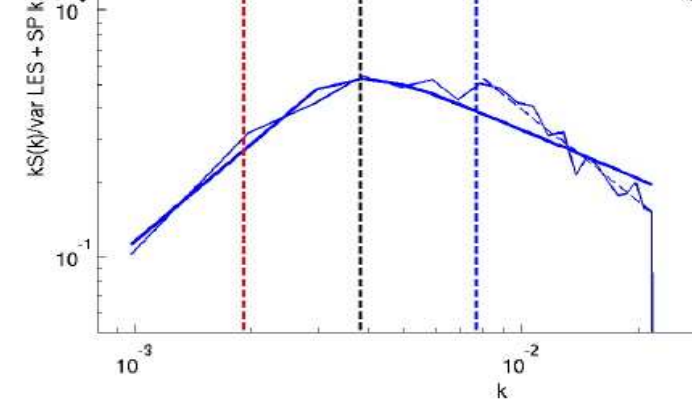
- Contributions (%)**
- **peak** : decreases from 16h30
  - **low** : increases from 16h30
  - **Inertial** : stays constant



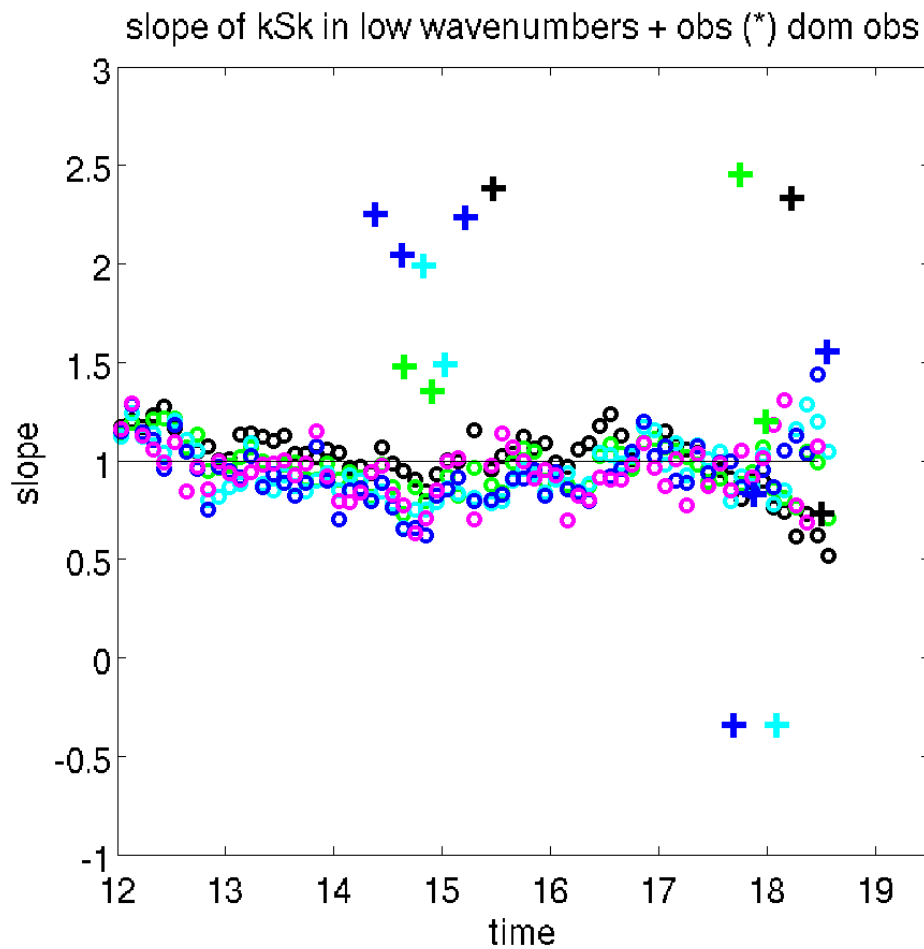
## II] How do the turbulent spectra evolve during the LAT?

- Results

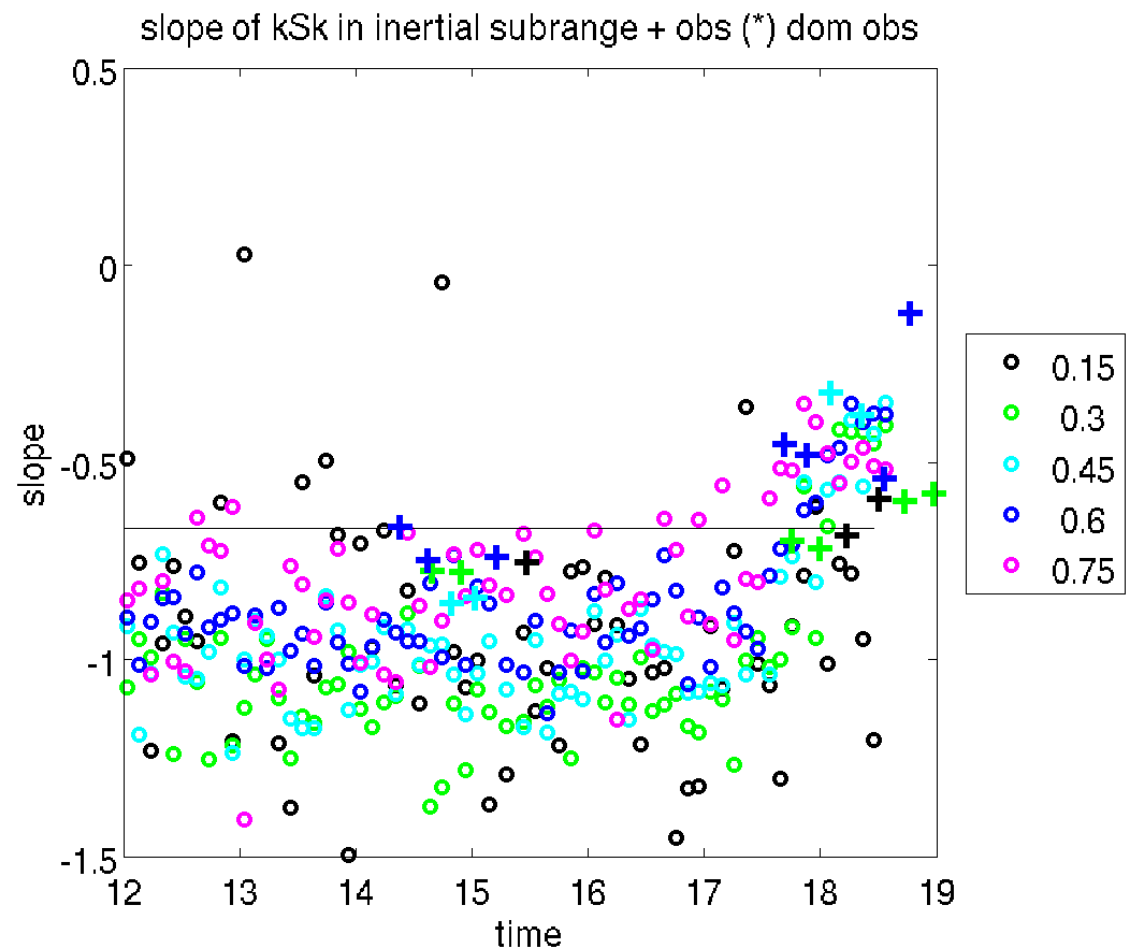
*LES for 20 June(o) + Obs of the 2 planes(+)*



### Evolution of the slopes in low wave numbers



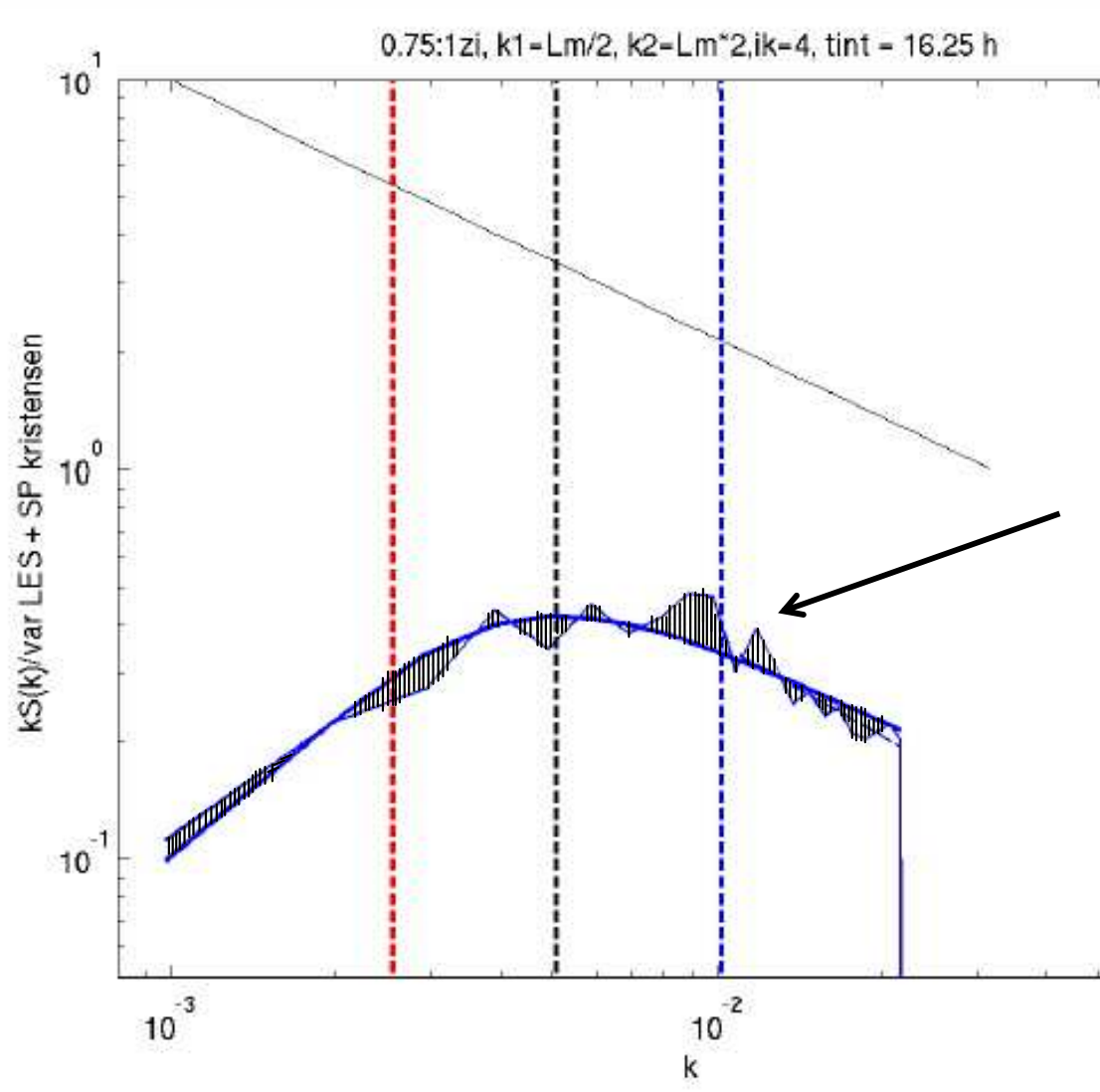
### Evolution of the slopes in the inertial subrange



## II] How do the turbulent spectra evolve during the LAT?

- Results

### Departure from the analytical model



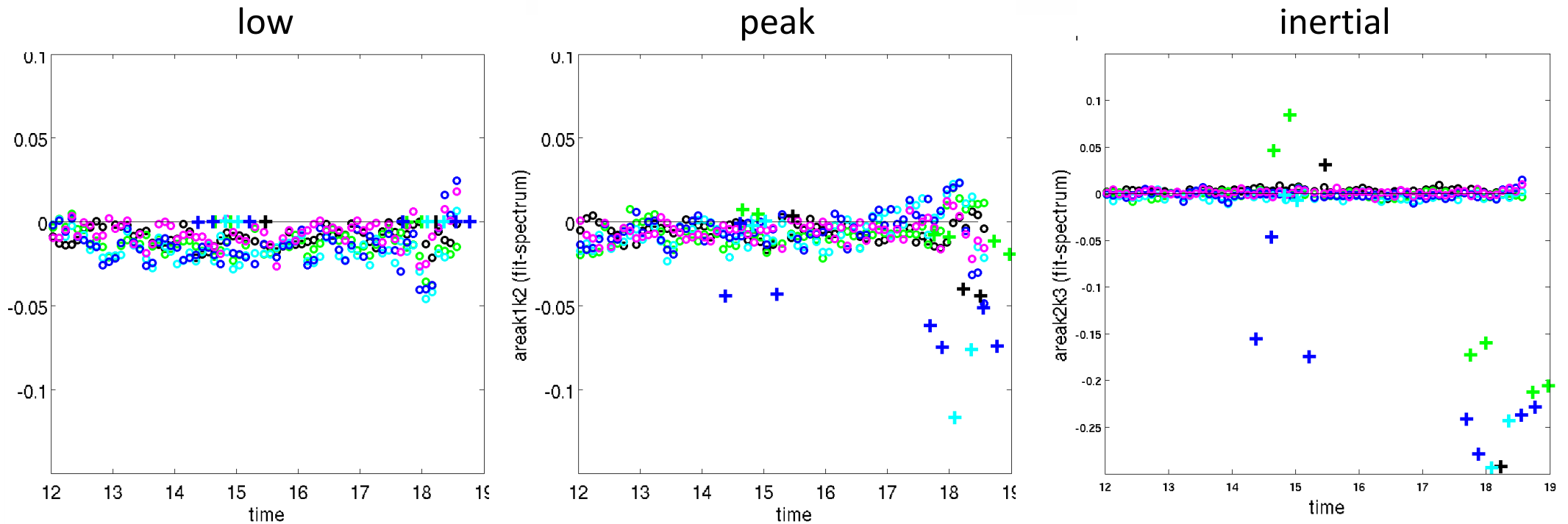
Integral over  $k$  of the difference between the fit and the LES spectra in the three domains

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- Integral over  $k$  of the difference between the fit and the LES spectra in the three domains



- Integral over  $k$  of the squared difference between the fit and the LES spectra
- Integral over  $k$  of the difference between  $\log(\text{fit})$  and  $\log(\text{LES})$  spectra
- Integral over  $k$  of the squared difference between  $\log(\text{fit})$  and  $\log(\text{LES})$  spectra

Ongoing work ...